## 2 - 16 A Test Study of γ Transitions from <sup>42,43</sup>Ti Nuclei

Xu Shiwei, He Jianjun, Hou Suqing, Yu Xiangqing, Zhang Ningtao Fang Yongde and Liu Minliang

The precise excitation energies of the proton-unbound states for the key nuclides along the rapid-proton process (so called rp-process) path play a very important role in nuclear astrophysics reaction rate calculations. This work reports the preliminary results of a test study of  $\gamma$  transitions from the <sup>42,43</sup> Ti nuclei.

The experiment was performed at the HI-13 Tandem Accelerator at China Institute of Atomic Energy (CIAE). A 35 MeV  $^{16}$ O beam bombarded an isotopically enrich  $^{28}$ Si metallic foil with a Pb backing (13.8 mg/cm² thick). The excited states in the  $^{42,43}$ Ti nuclei can be populated via the fusion-evaporation reaction of  $^{28}$ Si( $^{16}$ O, 2n) $^{42}$ Ti and  $^{28}$ Si( $^{16}$ O, n) $^{43}$ Ti, respectively. Once the excited states populated, the related  $\gamma$  transitions can be measured and identified by the X- $\gamma$ -t and  $\gamma$ - $\gamma$ -t coincidence technique. In the experiment, the  $\gamma$  rays detected with an array consisting of 14 Compton-suppressed HPGe detectors.

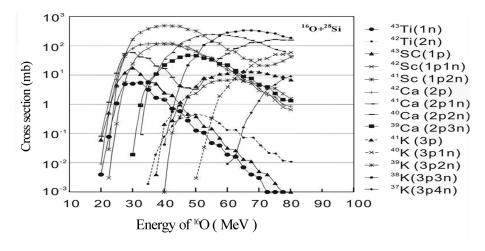


Fig. 1 Reaction cross sections for the <sup>16</sup> O+<sup>28</sup> Si colliding system calculated by an ALICE code<sup>[1]</sup>.

The reaction cross sections for the  $^{16}O+^{28}Si$  colliding system were calculated by an ALICE code  $^{[1]}$ , as shown in Fig. 1. However, none of  $\gamma$  transition was observed in the experiment because of the low cross sections for the  $^{42\cdot43}$ Ti production. The  $\gamma$  transitions from the excited states in  $^{42}Sc$ ,  $^{39}K$  and  $^{42}Ca$  were observed. We found that the cross section of the  $^{28}Si(^{16}O, 3p2n)^{39}K$  reaction was almost equal to that of the  $^{28}Si(^{16}O, 2p)^{42}Ca$  reaction, which is contrary to the theoretical calculations. The data analysis is still in progress.

## Reference

[1] M. Blann, H. K. Vonach, Phys. Rev., C28(1983)1475.